

Details of TPM - Discreteness Correlation

1.) Discreteness Correlation Function

$$\begin{cases} 1 = x_1, v_1, t \\ 2 = x_2, v_2, t \end{cases}$$

$$\tilde{f}(i) = \frac{1}{n} \sum_{i=1}^N \delta(x_i - x_i(t)) \delta(v_i - v_i(t))$$

$$\langle \rangle = \int dx_i \int dv_i n \langle f \rangle$$

$$\langle \rangle = \int dx_i \int dv_i n \langle f \rangle$$

$$\langle \tilde{f}(1) \tilde{f}(2) \rangle =$$

$$= \int dx_i \int dv_i n \left(\frac{1}{n} \sum_{i=1}^N \delta(x_i - x_i(t)) \delta(v_i - v_i(t)) \right) \left(\frac{1}{n} \sum_{j=1}^N \delta(x_j - x_j(t)) \delta(v_j - v_j(t)) \right)$$

$$= \int dx_i \int dv_i \frac{\langle f \rangle}{n} \sum_{\substack{i,j=1 \\ i \neq j}}^N \left[\delta(x_i - x_j(t)) \delta(x_j - x_i(t)) \delta(v_i - v_j(t)) \delta(v_j - v_i(t)) \right]$$

only $\neq 0$ if arguments interchangeable

$$= \int dx_i \int dv_i \frac{\langle f \rangle}{n} \sum_{\substack{i,j=1 \\ i \neq j}}^N \left[\delta(x_i - x_j) \delta(x_j - x_i) \delta(v_i - v_j) \delta(v_j - v_i) \right]$$

$$= \frac{\langle f \rangle}{n} \delta(x_1 - x_2) \delta(v_1 - v_2)$$

Now seek:

$$C(k, \omega) = \int dv_1 \int dv_2 \langle \tilde{f}(1) \tilde{f}(2) \rangle_{k, \omega}$$

$$= \int dv_1 \int dv_2 \int dx_- e^{-ikx_-} \int d\mathcal{T} e^{i\omega\mathcal{T}} \langle \tilde{f}(0) \tilde{f}(x_-, \mathcal{T}) \rangle$$

→ where we have used: $\left\{ \begin{array}{l} \text{homogeneity} \\ \text{stationarity} \end{array} \right.$

$$\rightarrow v_{\pm} = (v_2 \pm v_1)/2$$

so

$$\int dv_- \langle \tilde{f}(1) \tilde{f}(2) \rangle_{k, \omega} = \frac{\langle f \rangle}{n} \int e^{i\omega\mathcal{T}} e^{-ikx_-(\mathcal{T})} d\mathcal{T}$$

$x(\mathcal{T}) = v\mathcal{T} \rightarrow$ ballistic, u.p.o.

$$\int dK \langle \tilde{f}^2 \rangle_{k, \omega} = 2\pi \delta(\omega - kv) \frac{\langle f \rangle}{n}$$

\downarrow
 k, ω orbit propagator

\hookrightarrow weight function \rightarrow particle distribution

So
$$C(k, \omega) = \int dV \frac{\langle f \rangle}{n} 2\pi \delta(\omega - kv) \rightarrow \text{noise spectrum}$$

So, can write in 1D: (hereafter)

$$C(k, \omega) = \frac{2\pi}{n|k|v_{Te}} \langle \bar{f}(\omega/kv_{Te}) \rangle$$

↳ v_{Te} factor extracted

discreteness noise has Maxwellian Doppler spectrum...

and

$$\langle \hat{\phi}^2 \rangle_{n, \omega} = \left(\frac{4\pi n_0 e l}{k^2} \right)^2 \frac{1}{n_0} \frac{2\pi}{|k|v_{Te}} \frac{\langle \bar{f}(\omega/kv_{Te}) \rangle}{|E(k, \omega)|^2}$$

∴ have obtained thermal equilibrium spectrum.

Can re-write as:

$$\langle \phi^2 \rangle_{k, \omega} = n_0 \left(\frac{4\pi |e|}{k^2} \right)^2 \frac{2\pi}{k v_{Te}} \frac{\langle \bar{F}(\omega/k v_{Te}) \rangle}{|\epsilon(k, \omega)|^2}$$

↑ # particles
 ↑ Coulomb Spectrum
 ↑ Screening

→ Spectrum defined / determined by:

- e.g. particle emission spectrum $\sim e^{-\omega^2/k^2 v_{Te}^2}$

- collective response $\epsilon \approx 1 - \frac{\omega_p^2}{\omega^2} + i\epsilon_{IM}$

$(\omega > k v_{Te})$

$\epsilon \approx 1 + 1/k^2 N_D^2 + i\epsilon_{IM}$

$(\omega < k v_{Te})$

→ Can note:

1) collective response strongest at wave resonance $(k v_{Te} \sim \omega \sim \omega_{pe})$

⇒ expect peak at natural frequency.

2) $\omega \gg \omega_p$ ($v \gg \omega_p/k$) \Rightarrow noise source decouples from collective dynamics (i.e. plasma as vacuum)

$$\langle \hat{\phi}^2 \rangle_{k,\omega} \approx n_0 \left(\frac{4\pi|e|}{k^2} \right)^2 \frac{2\pi}{|k|v_{te}} e^{-\omega^2/k^2v_{te}^2}$$

3) $\omega \ll \omega_p \Rightarrow$ static limit (Debye screening)

$$\langle \hat{\phi}^2 \rangle_{k,\omega} \approx n_0 \left(\frac{4\pi|e|}{k^2} \right)^2 \frac{2\pi}{|k|v_{te}} e^{-\omega^2/k^2v_{te}^2} / \left(k^2 + 1/\lambda_D^2 \right)^2$$

So, can write Electric Field Spectrum :

$$\left| \frac{\hat{E}_{k,\omega}}{8\pi} \right|^2 = \frac{4\pi^2 n_0 |e|^2}{k^2 |k|} \left\{ \frac{F(\omega/kv_{te})}{\left(1 - \frac{\omega_p^2}{\omega^2} \right)^2 + \left(\frac{\pi \omega_p^2}{|k|k} F' \right)^2} \right\}$$

$$|E_r|^2 + |E_{EM}|^2$$

$F' \sim$ slope.

i.e. $F' = dF/d(\frac{\omega}{k})$

Now, to make contact with usual expectations:

i.e. $k_B T/2$ per degree-of-freedom

$$W_k = \int d\omega |\hat{E}_{k,\omega}|^2 / 8\pi$$

field energy per mode

Useful trick: Pole Approximation

$$\frac{1}{|\epsilon|^2} \approx \frac{1}{(\omega - \omega_n)^2 \left| \frac{d\epsilon}{d\omega} \right|^2 + |\epsilon_{IM}|^2} \rightarrow \text{sets width}$$

↳ real freq. sets location

$$\approx \frac{1}{|\epsilon_{IM}|} \left\{ \frac{|\epsilon_{IM}|}{(\omega - \omega_n)^2 \left| \frac{d\epsilon}{d\omega} \right|^2 + |\epsilon_{IM}|^2} \right\}$$

$$\approx \frac{1}{|\epsilon_{IM}|} \left| \left(\frac{d\epsilon}{d\omega} \right)_{\omega_n} \right|^{-1} \delta(\omega - \omega_n)$$

$$\frac{1}{|\epsilon|^2} \approx \delta(\omega - \omega_n) / |\epsilon_{IM}(\omega_n)| \left| \frac{d\epsilon}{d\omega} \right|_{\omega_n}$$

so, integrating in pole approximation:

$$W_k = m_e \omega_{pe} \frac{F}{2|k| |F'|}$$

$$\approx m_e \omega_{pe} \frac{F}{2|k|} \frac{1}{\frac{\omega_{pe}}{|k| v_{Te}} F} = \frac{T}{2} \checkmark$$

in accord with "T/2 per degree" intuition

→ if $k \lambda_D \gg 1$ ($\epsilon \rightarrow 1$ limit)

$$W_k \approx \frac{T}{2} \frac{1}{k^2 \lambda_D^2} \rightarrow \text{strong cut-off beyond } \lambda_D.$$

So, for total energy density:

$$\left\langle \frac{E^2}{8\pi} \right\rangle = \int d\underline{k} W_{\underline{k}} = \int \frac{d\underline{k}}{(2\pi)^3} \frac{T/2}{1+k^2 \lambda_D^2}$$

$$\sim \left(\frac{k_B T}{2} \right) k_{\max}^3$$

$$\sim \left(\frac{n k_B T}{2} \right) \frac{1}{n \lambda_D^3}$$

so

$$\left\langle \frac{E^2}{8\pi} \right\rangle \sim \left(\frac{n k_B T}{2} \right) \frac{1}{n \lambda_D^3}$$

$$\sim (\text{Particle Kinetic Energy Density}) / n \lambda_D^3$$

↓
particles
in Debye sphere

⇒

$$(FED) \sim \frac{(PKED)}{n \lambda_D^3}$$

$1/n \lambda_D^3 \sim$ discreteness factor

→ To connect formally, to fluctuation-dissipation theorem:

$$\text{Note: } \epsilon_{IM} = \frac{-\omega_p^2 \pi}{k|k|} \frac{\partial \langle f \rangle}{\partial V} \Big|_{\omega/k}$$

$$= \frac{2\pi \omega}{k^2 v_{te}^2 |k| v_{te}} \frac{\omega_p^2}{\epsilon} \langle \bar{f}(\omega/k) \rangle, \quad \text{for Maxwellian } \langle f \rangle$$

$$\text{so } \langle \bar{f}(\omega/k) \rangle = k^2 v_{te}^2 |k| v_{te} \epsilon_{IM} / 2\pi \omega \omega_p^2$$

and have:

$$\langle \hat{\phi}^2 \rangle_{k,\omega} = \frac{2\pi T}{|k| v_{te}} \left(\frac{4\pi |e|}{k^2} \right)^2 \frac{\langle \bar{f}(\omega/k) \rangle}{|\epsilon(k,\omega)|^2}$$

so, plugging in:

$$\langle \hat{\phi}^2 \rangle_{k,\omega} = \frac{8\pi T}{k^2 \omega} \frac{\text{Im} \epsilon}{|\epsilon|^2}$$

and

$$\left\langle \frac{\hat{E}^2}{8\pi} \right\rangle_{k,\omega} = \frac{T}{\omega} \frac{\text{Im} \epsilon}{|\epsilon|^2}$$

↔ { Fluctuation-
Dissipation
Theorem

(restates form of spectrum)

→ relates thermal fluctuations to dissipation in collective modes ($\text{Im} \epsilon$)

→ obviously consistent (by construction) with physical picture of thermal equilibrium

b.) Some general comments:

→ Key element of T.P.M. is use of linear $F_{k,\omega}^c$ or, equivalently, unperturbed orbits,

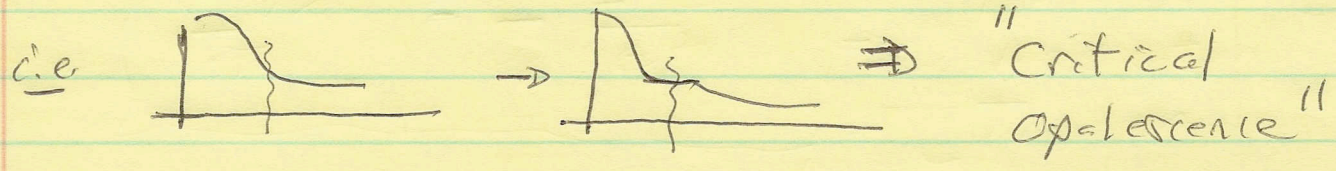
- This assumes small fluctuation levels, so stochastic deflection is 'weak'

d.e. $\underline{x}(t) = \underline{x}(0) + \underline{v}t + \delta(\underline{x}(t))$
deflection

How weak? ⇔ take care $T_{ac} < T_T$ condition

but - $\langle \hat{\phi}^2 \rangle_k \sim () \frac{F(\omega_{pe}/k)}{|F'(\omega_{pe}/k)|}$

∴ Fluctuations diverge as $F' \rightarrow 0$, from below



Note $F' > 0$ not necessary ⇔ theory

fails for stable plasma....., approaching marginality.

→ As fluctuations grow, linearizations fail

∴ must renormalize!

→ particle propagator

$$c^i / \omega - kv \rightarrow c^i / \omega - kv + \sum_{\text{self energy,}}^{\text{decoration rate}}$$

→ mode propagator / response

$$I/E \rightarrow \frac{1}{\left[\underbrace{\omega - (\omega_r + \delta\omega_r)}_{\text{nonlinear frequency shift}} \frac{\partial \epsilon}{\partial \omega} + i \left(\underbrace{\epsilon_{IH}^L + \epsilon_{IH}^{NL}}_{\substack{\text{nonlinear} \\ \text{dissipation} \\ (\omega-\omega \text{ interaction}) \\ (\omega \rightarrow p \text{ interaction}) \\ \Rightarrow \gamma_{NL}}} \right) \right]}$$

(recall NL oscillator, driven)

Calculating all this is aim of plasma turbulence theory.....